# Correlation of Optical Image Sensor Noise in Space with Trapped Proton Flux

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Michael Zambrana

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## 14. ABSTRACT

MPTB particle environment measurements are being correlated with noise signals in a CMOS visible sensor to model interactions of particles and pixel sensor elements.

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## Summary & Introduction

MPTB is flying in a highly elliptical HEO orbit that passes through the proton belts. The CMOS active pixel optical sensor in a related experiment (called TX) shows increased noise levels during proton belt passages so noise signals above background can be attributed to the deposition and collection of charge from protons. Another related experiment flying with MPTB is a particle detector which measures, among other items, proton flux. This work reports correlations of pixel noise and proton flux in order to quantify the response of the sensor elements to a measured proton spectrum. This is done through modeling of radiation transport to the sensor elements with a detailed knowledge of the design and fabrication of the sensor. These results will then be used to predict the response of more modern optical sensors to proton environments.

The MPTB orbit is very close to 12 hours in duration; the perigee is about 1000 km, while apogee is beyond geo-synchronous. Nearly 9 hours of the orbit are in deep space, with the remaining 3.5 hours comprised of a rapid drop to perigee and climb back to deep space. Thus the orbital path transits directly through the proton belts twice each orbit. Details of the orbit can be found in Dyer et.al.[1]

An instrument flying with MPTB is being utilized to provide measurements of the radiation environment at the same time the measurements are being made on the sensor. It is the Aerospace DSU (<u>DoSimetry Unit</u>) instrument. The DSU measures proton fluxes in four energy regions and dose behind several shield thicknesses over a hemispherical geometry. For this work, the proton fluxes with energies greater than 6.5, 15 and 25 MeV are being analyzed. The DSU instrument is described in detail in references[2-7]. Figure 1 includes DSU proton flux data for one orbit segment near perigee on 30 September 2002. Data is taken approximately each 15 seconds.

Measurements were made using the TX instrument during proton belt passes. The signals analyzed were due to inherent noise in the sensor and related electronics and interactions with the radiation environment. Sequential reads row-by-row are performed and this data is what is used in this analysis. Data in this summary represent arbitrary noise signals but in the full paper this will be converted to electron-hole pairs using the overall gain and sensitivity of instrument.

The visible sensor data described in this summary in the TX instrument were acquired using a Rockwell TCM2620V monolithic silicon sensor array fabricated in a 0.6 micrometer CMOS process. The sensor has elements with a 40 micrometer pitch, a 30% active area, and a depletion thickness of 12 micrometers. The format is 256 x 256 pixels with multiplexed snapshot mode Capacitive Trans-Impedance Amplifier (CTIA) readout. Unit cell area was divided between the CTIA elements, including a 0.4 pF capacitor, and implanted n-on-p diode structures for optical detection. The sensor includes on-chip clock and bias generators as well as column and row buffering, and four high-speed class A-B amplifier outputs. With operation at room temperature, the read noise was typically 30 electrons with a 3-volt maximum signal. Further descriptions of the n-on-p detector optical characteristics and circuit performance will be provided in the full paper.

Proton testing was performed on samples of the TCM2620V array, and the device was found to be insensitive to proton induced latchup. Also, on-orbit flight data will be compared with the measured response and modeling with 60 MeV protons incident with a flux of 8.54x10<sup>6</sup> p/cm<sup>2</sup>/s at a 15 Hz readout rate. Additional details of the laboratory test detail and measurements will be provided in the full paper.

## Modeling

The host vehicle and the TX experiment were modeled with the Air Force Research Laboratory radiation modeling code SVC[8]. The effective average shielding, in terms of 528 aluminum equivalent rays, was determined to be about 650 mils of aluminum about the TX sensor. The distribution shows 177 of these are 1000 mils or less. Figure 2 shows the distribution of these rays. Although the DSU provides the trapped proton spectra for protons above 25 MeV, only incident protons of greater than 65 MeV can reach the sensor. The AP-8min flux map was used to generate the high-energy portion of the incident proton flux at the sensor location. Figure 3 shows the incident and penetrating flux, integral in energy, at the sensor location for the first proton peak of Figure 1. Clearly the sensor shielding results in a hard proton energy spectrum at the sensor. A more detailed shielding analysis to consider the sensor orientation and possible directional effects from the shielding will be performed for the presentation and paper. Proton spectra at four shield thicknesses, 460 (minimum), 650, 1000 and 2000 mils equivalent Al will be used for the analysis.

In addition to the DSU proton flux measurements, Figure 4 also shows the largest pixel noise signals for the same time period. Rows of 170 pixels, a segment of the 256 cells per row, are sampled four at a time approximately each 1.8 seconds and the figure shows the largest value among the 170 in arbitrary units. The average noise level inherent in the sensor is 4.4.

Figure 5 shows dark frames, no light input, of on-orbit taken during two portions of an orbit. The top figure a shows an image taken outside of the proton belts and the bottom figure b shows an image taken during a proton belt passage. This data is histogrammed and converted to charge carrier for the analysis. This will then be used in the modeling.

In the full paper we will provide detailed predictions of the expected array response relative to the trapped proton environment after transport through surrounding structures. Due to the confinement of the n-on-p diode to the epitaxial layer, we model the charge collection as a combination of drift with some contribution from diffusion within the 12 micrometer epitaxial layer. Assuming a 12 micrometer depletion depth, the resulting charge due to a strike by a trapped proton is typically several thousand electrons, with the charge yield depending on proton energy, trajectory, and to some extent the carrier lifetime in the epitaxial layer. Also, since the implanted diode occupies only a portion of the unit cell, we do not expect charge sharing between adjacent pixels due to diffusion. The paper will

describe both in measured array response and model predictions, that adjacent pixels are affected only when proton trajectories result in transport through neighboring detector elements.

Our approach to modeling the interactions between transported protons and the pixel level response is based on a computational tool that was described by Pickel [7]. Modifications from the cited application to infrared detectors include the incorporation of the Si detector material, and geometries appropriate to monolithic detection in a CMOS implanted diode. The model employs a combination of analytic and Monte Carlo techniques to track ion deposited charge as it is collected via both drift and diffusion in a two dimensional array. Structural dependent spectral characteristics of the transported environment with detailed descriptions of trajectories are incorporated in the modeling approach. Further details of the model and comparisons between modeled and measured results will be presented in the full paper.

## **Conclusions**

This work will show the correlations in space between measured particle flux and spectrum and the noise interactions of a CMOS image sensor pixel. This will be done using ray-tracing modeling including a detailed analysis of the sensor design.

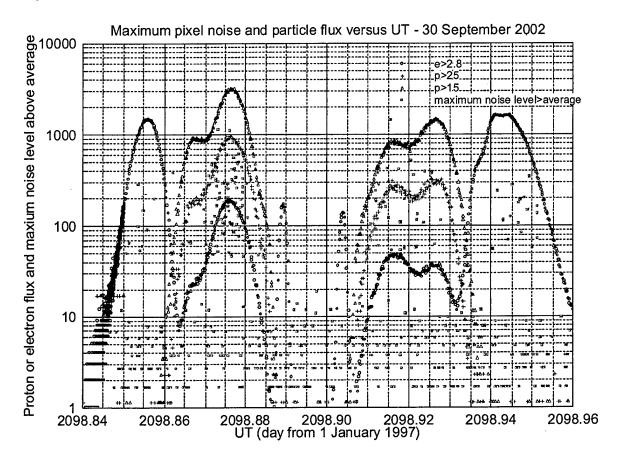


Figure 1: Maximum row pixel noise signal above average & proton & electron flux vs UT - 30 September 2002.

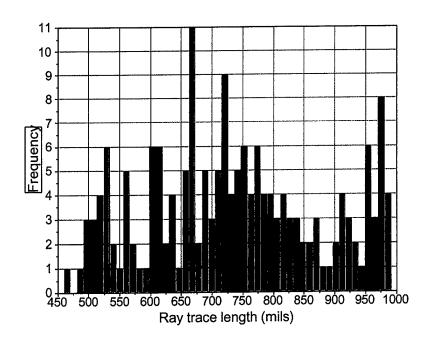


Figure 2: Length distribution histogram of rays less than 1000 mils.

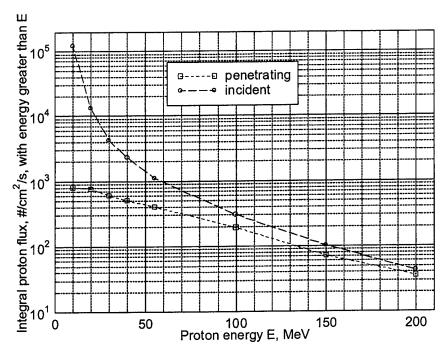


Figure 3 – Calculated incident proton spectra incident and penetrating (650 mils) the TX instrument striking the sensor.

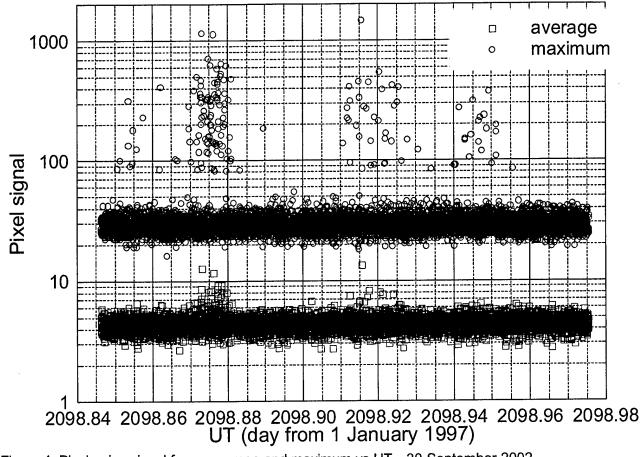
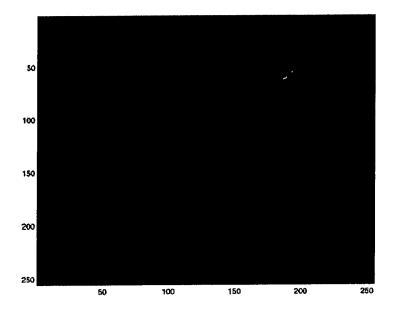


Figure 4: Pixel noise signal frame average and maximum vs UT - 30 September 2002



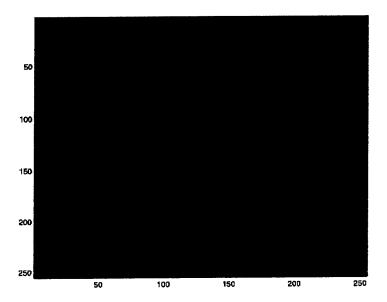


Figure 5a & 5b: The two figures above show output from the TX sensor while dark measurements were made outside of the proton belts (top, a) and during a proton belt traversal (bottom, b).

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